Weak solutions of conservation laws and energy/entropy conservation

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Introduction: the principle of conservation of energy for classical solutions

Let us first focus our attention on the incompressible Euler system

$$\partial_t u + \operatorname{div}(u \otimes u) + \nabla p = 0,$$

 $\operatorname{div} u = 0,$

If u is a classical solution, then multiplying the balance equation by u we obtain

$$\frac{1}{2}\partial_t |u|^2 + \frac{1}{2}u \cdot \nabla |u|^2 + u \cdot \nabla p = 0.$$

Integrating the last equality over the space domain Ω yields

$$\frac{\mathrm{d}}{\mathrm{d}t} \int_{\Omega} \frac{1}{2} |u(x,t)|^2 \, \mathrm{d}x = 0.$$

Consequently, integrating over time in (0, t), gives

$$\int_{\Omega} \frac{1}{2} |u(x,t)|^2 dx = \int_{\Omega} \frac{1}{2} |u(x,0)| dx.$$

Weak solutions

However, if u is a weak solution, then

$$\int_{\Omega} \frac{1}{2} |u(x,t)|^2 \ \mathrm{d}x = \int_{\Omega} \frac{1}{2} |u(x,0)| \ \mathrm{d}x.$$

might not hold. Technically, the problem is that u might not be regular enough to justify integration by parts in the above derivation.

Motivated by the laws of turbulence Onsager postulated that there is a critical regularity for a weak solution to be a conservative one:

Conjecture, 1949

Let u be a weak solution of incompressible Euler system

- If $u \in C^{\alpha}$ with $\alpha > \frac{1}{3}$, then the energy is conserved.
- For any $\alpha < \frac{1}{3}$ there exists a weak solution $u \in C^{\alpha}$ which does not conserve the energy.

Onsager conjecture for incompressible Euler system

Weak solutions of the incompressible Euler equations which do not conserve energy were constructed:

- Scheffer '93, Shnirelman '97 constructed examples of weak solutions in $L^2(\mathbb{R}^2 \times \mathbb{R})$ compactly supported in space and time
- De Lellis and Székelyhidi 2010 showed how to construct weak solutions for given energy profile

Still incompressible case

Onsager conjecture:

If weak solution v has $C^{0,\alpha}$ (for $\alpha > \frac{1}{3}$) regularity then it conserves energy. In the opposite case it may not conserve energy.

- The first part of this assertion was proved in
 - P. Constantin, W. E, and E. S. Titi. Onsager's conjecture on the energy conservation for solutions of Euler's equation.
 Comm. Math. Phys., 1994
 - G. L. Eyink. Energy dissipation without viscosity in ideal hydrodynamics. I. Fourier analysis and local energy transfer. Phys. D, 1994
 - A. Cheskidov, P. Constantin, S. Friedlander, and R. Shvydkoy. Energy conservation and Onsager's conjecture for the Euler equations. Nonlinearity, 2008

Besov spaces

The elements of Besov space $B_p^{\alpha,\infty}(\Omega)$, where $\Omega=(0,T)\times\mathbb{T}^d$ or $\Omega=\mathbb{T}^d$ are functions w for which the norm

$$\|w\|_{B^{\alpha,\infty}_p(\Omega)}:=\|w\|_{L^p(\Omega)}+\sup_{\xi\in\Omega}\frac{\|w(\cdot+\xi)-w\|_{L^p(\Omega\cap(\Omega-\xi))}}{|\xi|^\alpha}$$

is finite (here $\Omega - \xi = \{x - \xi : x \in \Omega\}$).

It is then easy to check that the definition of the Besov spaces implies

$$\|w^{\epsilon} - w\|_{L^{p}(\Omega)} \le C\epsilon^{\alpha} \|w\|_{B_{p}^{\alpha,\infty}(\Omega)}$$

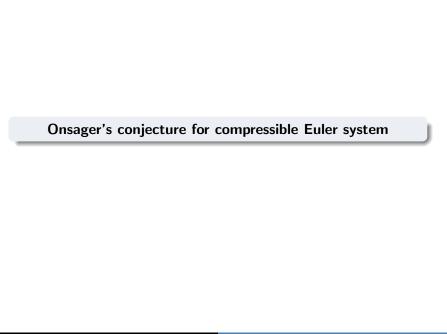
and

$$\|\nabla w^{\epsilon}\|_{L^{p}(\Omega)} \leq C\epsilon^{\alpha-1}\|w\|_{B_{p}^{\alpha,\infty}(\Omega)}.$$

Idea of the proof: P. Constantin, W. E, and E. S. Titi. Onsager's conjecture on the energy conservation for solutions of Euler's equation. Comm. Math. Phys., 1994

- ullet take as the test function doubly mollified solution $(v^\epsilon)^\epsilon$
- problem: estimate term $\int_{\mathbb{T}^d} \operatorname{Tr}(v \otimes v)^{\epsilon} \cdot \nabla v^{\epsilon} dx$
- use the identity:

$$(v \otimes v)^{\epsilon} = v^{\epsilon} \otimes v^{\epsilon} + r_{\epsilon}(v, v) - (v - v^{\epsilon}) \otimes (v - v^{\epsilon})$$
 where $\|r_{\epsilon}(v, v)\|_{L^{3/2}} \leq C \epsilon^{2\alpha} \|v\|_{\mathcal{B}^{\alpha, \infty}_{p}}^{2}$



Compressible Euler system

We consider now the isentropic Euler equations,

$$\partial_{t}(\rho u) + \operatorname{div}(\rho u \otimes u) + \nabla p(\rho) = 0,$$

$$\partial_{t}\rho + \operatorname{div}(\rho u) = 0.$$
 (1)

We will use the notation for the so-called pressure potential defined as

$$P(\rho) = \rho \int_1^\rho \frac{p(r)}{r^2} dr.$$

Theorem (Feireisl, Gwiazda, Ś.-G., Wiedemann, ARMA 2017)

Let ϱ , u be a solution of (1) in the sense of distributions. Assume

$$u \in B_3^{\alpha,\infty}((0,T) \times \mathbb{T}^d), \ \varrho, \varrho u \in B_3^{\beta,\infty}((0,T) \times \mathbb{T}^d), 0 \leq \underline{\varrho} \leq \varrho \leq \overline{\varrho}$$

for some constants ϱ , $\overline{\varrho}$, and $0 \le \alpha, \beta \le 1$ such that

$$\beta > \max\left\{1 - 2\alpha; \frac{1 - \alpha}{2}\right\}. \tag{2}$$

Assume further that $p \in C^2[\varrho, \overline{\varrho}]$, and, in addition

$$p'(0) = 0$$
 as soon as $\varrho = 0$.

Then the energy is locally conserved in the sense of distributions on $(0, T) \times \Omega$, i.e.

$$\partial_t \left(\frac{1}{2} \varrho |u|^2 + P(\varrho) \right) + \operatorname{div} \left[\left(\frac{1}{2} \varrho |u|^2 + p(\varrho) + P(\varrho) \right) u \right] = 0.$$

Sharpness of assumptions

Shocks provide examples that show that our assumptions are sharp:

- A shock solution dissipates energy, but ρ and u are in $BV \cap L^{\infty}$, which embeds into $B_3^{1/3,\infty}$.
- Hence such a solution satisfies (2) with equality but fails to satisfy the conclusion.

Time regularity

The hypothesis on temporal regularity can be relaxed provided

$$\varrho > 0$$

Indeed, in this case $\frac{(\varrho u)^{\epsilon}}{\varrho^{\epsilon}}$ can be used as a test function in the momentum equation, cf.

T. M. Leslie and R. Shvydkoy. The energy balance relation for weak solutions of the density-dependent Navier- Stokes equations. JDE, 2016.

Some references



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General conservation laws

- It is easy to notice similarities in the statements regarding sufficient regularity conditions guaranteeing energy/entropy conservation for various systems of equations of fluid dynamics.
- Especially the differentiability exponent of $\frac{1}{3}$ is a recurring condition.
- One might therefore anticipate that a general statement could be made, which would cover all the above examples and more.
 Indeed, consider a general conservation law of the form

$$\operatorname{div}_X(G(U(X))) = 0.$$

We consider the conservation law of the form

$$\operatorname{div}_X(G(U(X))) = 0. \tag{3}$$

Here $U: \mathcal{X} \to \mathcal{O}$ is an unknown and $G: \mathcal{O} \to \mathbb{M}^{n \times (d+1)}$ is a given, where \mathcal{X} is an open subset of \mathbb{R}^{d+1} or $\mathbb{T}^3 \times \mathbb{R}$ and the set \mathcal{O} is open in \mathbb{R}^n . It is easy to see that any classical solution to (3) satisfies also

$$\operatorname{div}_X(Q(U(X))) = 0, \tag{4}$$

where $Q:\mathcal{O} o \mathbb{R}^{\mathbf{s} imes (d+1)}$ is a smooth function such that

$$D_U Q_j(U) = \mathfrak{B}(U) D_U G_j(U)$$
, for all $U \in \mathcal{O}, j \in 0, \dots, k$, (5)

for some smooth function $\mathfrak{B}:\mathcal{O}\to\mathbb{M}^{s\times n}$. The function Q is called a *companion* of G and equation (4) is called a *companion law* of the conservation law (3).

How much regularity of a weak solution is required so that it also satisfies the companion law?

Theorem

Let $U \in B_3^{\alpha,\infty}(\mathcal{X};\mathcal{O})$ be a weak solution of (3) with $\alpha > \frac{1}{3}$. Assume that $G \in \mathrm{C}^2(\mathcal{O}; \mathbb{M}^{n \times (d+1)})$ is endowed with a companion law with flux $Q \in \mathrm{C}(\mathcal{O}; \mathbb{M}^{s \times (d+1)})$ for which there exists $\mathfrak{B} \in \mathrm{C}^1(\mathcal{O}; \mathbb{M}^{s \times n})$ related through identity (5) and the essential image of U is compact in \mathcal{O} .

Then U is a weak solution of the companion law (4) with the flux Q.



P. Gwiazda, M. Michálek and A. Świerczewska-Gwiazda.

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Remarks

- the generality of the above theorem is achieved at the expense of optimality of the assumptions.
- However given additional information on the structure of the problem at hand one might be able to relax some of these assumptions.
- the theorem provides for instance a conservation of energy result for the system of polyconvex elastodynamics, compressible hydrodynamics et al.
- T. Debiec, P. Gwiazda, and A. Świerczewska-Gwiazda, A tribute to conservation of energy for weak solutions arXiv:1707.09794, 2017.

Thank you for your attention