# <span id="page-0-0"></span>Weak solutions of conservation laws and energy/entropy conservation

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Roztoky, July 3rd, 2017

# Introduction: the principle of conservation of energy for classical solutions

Let us first focus our attention on the incompressible Euler system

$$
\partial_t u + \operatorname{div}(u \otimes u) + \nabla p = 0,
$$
  
div  $u = 0$ ,

If  $u$  is a classical solution, then multiplying the balance equation by u we obtain

$$
\frac{1}{2}\partial_t|u|^2+\frac{1}{2}u\cdot\nabla|u|^2+u\cdot\nabla p=0.
$$

Integrating the last equality over the space domain  $\Omega$  yields

$$
\frac{\mathrm{d}}{\mathrm{d}t}\int_{\Omega}\frac{1}{2}|u(x,t)|^2\,\mathrm{d}x=0.
$$

Consequently, integrating over time in  $(0, t)$ , gives

$$
\int_{\Omega} \frac{1}{2} |u(x, t)|^2 dx = \int_{\Omega} \frac{1}{2} |u(x, 0)| dx.
$$

# Weak solutions

However, if  $u$  is a weak solution, then

$$
\int_{\Omega} \frac{1}{2} |u(x, t)|^2 dx = \int_{\Omega} \frac{1}{2} |u(x, 0)| dx.
$$

might not hold. Technically, the problem is that  $u$  might not be regular enough to justify integration by parts in the above derivation.

Motivated by the laws of turbulence Onsager postulated that there is a critical regularity for a weak solution to be a conservative one:

## Conjecture, 1949

Let  $u$  be a weak solution of incompressible Euler system

- If  $u \in \mathbb{C}^{\alpha}$  with  $\alpha > \frac{1}{3}$ , then the energy is conserved.
- For any  $\alpha < \frac{1}{3}$  there exists a weak solution  $u \in C^{\alpha}$  which does not conserve the energy.

Weak solutions of the incompressible Euler equations which do not conserve energy were constructed:

- Scheffer '93, Shnirelman '97 constructed examples of weak solutions in  $L^2(\mathbb{R}^2\times\mathbb{R})$  compactly supported in space and time
- De Lellis and Székelyhidi 2010 showed how to construct weak solutions for given energy profile

## Onsager conjecture:

If weak solution  $v$  has  $C^{0,\alpha}$  (for  $\alpha > \frac{1}{3}$ ) regularity then it conserves energy. In the opposite case it may not conserve energy.

- The first part of this assertion was proved in
	- P. Constantin, W. E, and E. S. Titi. Onsager's conjecture on the energy conservation for solutions of Euler's equation. Comm. Math. Phys., 1994
	- G. L. Eyink. Energy dissipation without viscosity in ideal hydrodynamics. I. Fourier analysis and local energy transfer. Phys. D, 1994
	- A. Cheskidov, P. Constantin, S. Friedlander, and R. Shvydkoy. Energy conservation and Onsager's conjecture for the Euler equations. Nonlinearity, 2008

The elements of Besov space  $B^{\alpha,\infty}_p(\Omega)$ , where  $\Omega=(0,\,T)\times\mathbb{T}^d$  or  $\Omega=\mathbb{T}^d$  are functions  $w$  for which the norm

$$
\|w\|_{B^{\alpha,\infty}_{p}(\Omega)} := \|w\|_{L^{p}(\Omega)} + \sup_{\xi \in \Omega} \frac{\|w(\cdot+\xi) - w\|_{L^{p}(\Omega \cap (\Omega-\xi))}}{|\xi|^{\alpha}}
$$

is finite (here  $\Omega - \xi = \{x - \xi : x \in \Omega\}$ ). It is then easy to check that the definition of the Besov spaces implies

$$
\|w^\epsilon - w\|_{L^p(\Omega)} \leq C\epsilon^\alpha \|w\|_{B^\alpha_p,^\infty(\Omega)}
$$

and

$$
\|\nabla w^\epsilon\|_{L^p(\Omega)}\leq C\epsilon^{\alpha-1}\|w\|_{B^{\alpha,\infty}_p(\Omega)}.
$$

Idea of the proof: P. Constantin, W. E, and E. S. Titi. Onsager's conjecture on the energy conservation for solutions of Euler's equation. Comm. Math. Phys., 1994

- take as the test function doubly mollified solution  $(\nu^{\epsilon})^{\epsilon}$
- problem: estimate term  $\int_{\mathbb{T}^d} \text{Tr} (\textbf{\textit{v}} \otimes \textbf{\textit{v}})^\epsilon \cdot \nabla \textbf{\textit{v}}^\epsilon d \textbf{\textit{x}}$
- o use the identity:

$$
(v \otimes v)^{\epsilon} = v^{\epsilon} \otimes v^{\epsilon} + r_{\epsilon}(v, v) - (v - v^{\epsilon}) \otimes (v - v^{\epsilon}) \text{ where}
$$
  

$$
||r_{\epsilon}(v, v)||_{L^{3/2}} \leq C\epsilon^{2\alpha} ||v||^{2}_{B^{\alpha, \infty}_{\rho}}
$$

# Onsager's conjecture for compressible Euler system

We consider now the isentropic Euler equations,

<span id="page-8-0"></span>
$$
\partial_t(\rho u) + \operatorname{div}(\rho u \otimes u) + \nabla p(\rho) = 0, \n\partial_t \rho + \operatorname{div}(\rho u) = 0.
$$
\n(1)

We will use the notation for the so-called pressure potential defined as

$$
P(\rho) = \rho \int_1^{\rho} \frac{p(r)}{r^2} dr.
$$

## Theorem (Feireisl, Gwiazda, S.-G., Wiedemann, ARMA 2017) ´

Let  $\varrho$ , u be a solution of [\(1\)](#page-8-0) in the sense of distributions. Assume  $u \in B_3^{\alpha,\infty}$  $\mathcal{L}_3^{\alpha,\infty}((0, T) \times \mathbb{T}^d), \varrho, \varrho u \in \mathcal{B}_3^{\beta,\infty}$  $\mathbb{C}^{ \beta, \infty }_{ 3 } ((0,\,T) \! \times \mathbb{T}^{d} ), 0 \leq \underline{\varrho} \leq \varrho \leq \overline{\varrho}$ 

for some constants  $\varrho$ ,  $\overline{\varrho}$ , and  $0 \leq \alpha, \beta \leq 1$  such that

<span id="page-9-0"></span>
$$
\beta > \max\left\{1 - 2\alpha; \frac{1 - \alpha}{2}\right\}.
$$
 (2)

Assume further that  $p\in\mathcal{C}^2[\varrho,\overline{\varrho}]$ , and, in addition

 $p'(0) = 0$  as soon as  $\varrho = 0$ .

Then the energy is locally conserved in the sense of distributions on  $(0, T) \times \Omega$ , i.e.

$$
\partial_t \left( \frac{1}{2} \varrho |u|^2 + P(\varrho) \right) + \text{div} \left[ \left( \frac{1}{2} \varrho |u|^2 + p(\varrho) + P(\varrho) \right) u \right] = 0.
$$

Agnieszka Swierczewska | [Energy/entropy conservation](#page-0-0)

Shocks provide examples that show that our assumptions are sharp:

- A shock solution dissipates energy, but  $\rho$  and  $u$  are in  $BV \cap L^{\infty}$ , which embeds into  $B_3^{1/3,\infty}$  $3^{1/3, \infty}$ .
- Hence such a solution satisfies [\(2\)](#page-9-0) with equality but fails to satisfy the conclusion.

The hypothesis on temporal regularity can be relaxed provided

$$
\underline{\varrho}>0
$$

Indeed, in this case  $\frac{(\varrho u)^{\epsilon}}{a^{\epsilon}}$  $\frac{\partial u_j}{\partial e^{\epsilon}}$  can be used as a test function in the momentum equation, cf.

T. M. Leslie and R. Shvydkoy. The energy balance relation for weak solutions of the density-dependent Navier- Stokes equations. JDE, 2016.



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Remarks on singularities, dimension and energy dissipation for ideal hydrodynamics and MHD. Comm. Math. Phys., 1997.



#### E. Kang and J. Lee.

Remarks on the magnetic helicity and energy conservation for ideal magneto-hydrodynamics. Nonlinearity, 2007.



#### R. Shvydkoy.

On the energy of inviscid singular flows. J. Math. Anal. Appl., 2009.



## R. Shvydkoy.

Lectures on the Onsager conjecture. Discrete Contin. Dyn. Syst. Ser. S, 2010.

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Regularity and Energy Conservation for the Compressible Euler Equations. Arch. Rational Mech. Anal., 2017.

S.

### J. Robinson, J. L. Rodrigo, J. W. D. Skipper.

Energy conservation in the 3D Euler equations on  $\mathcal{T}^2\times\mathbb{R}_+$  arXiv:, (1611.00181), 2017



# 量 C. Yu.

Energy conservation for the weak solutions of the compressible Navier–Stokes equations. Arch. Rational Mech. Anal., 2017.

#### T. M. Leslie and R. Shvydkoy.

The energy balance relation for weak solutions of the density-dependent Navier-Stokes equations. J. Differential Equations, 2016.



### T. D. Drivas and G. L. Eyink.

An Onsager singularity theorem for turbulent solutions of compressible Euler equations. to appear in Comm. in Math. Physics, 2017.



#### C. Bardos, E. Titi.

Onsager's Conjecture for the Incompressible Euler Equations in Bounded Domains. arXiv, (1707.03115), 2017.

- It is easy to notice similarities in the statements regarding sufficient regularity conditions guaranteeing energy/entropy conservation for various systems of equations of fluid dynamics.
- Especially the differentiability exponent of  $\frac{1}{3}$  is a recurring condition.
- One might therefore anticipate that a general statement could be made, which would cover all the above examples and more. Indeed, consider a general conservation law of the form

<span id="page-15-0"></span> $div_X(G(U(X))) = 0.$ 

We consider the conservation law of the form

$$
\operatorname{div}_X(G(U(X)))=0. \hspace{1.5cm} (3)
$$

Here  $U: \mathcal{X} \to \mathcal{O}$  is an unknown and  $G: \mathcal{O} \to \mathbb{M}^{n \times (d+1)}$  is a given, where  $\mathcal X$  is an open subset of  $\mathbb R^{d+1}$  or  $\mathbb T^3 \times \mathbb R$  and the set  $\mathcal O$ is open in  $\mathbb{R}^n$ . It is easy to see that any classical solution to  $(3)$ satisfies also

<span id="page-16-0"></span>
$$
\text{div}_X(Q(U(X)))=0,\qquad \qquad (4)
$$

where  $Q: \mathcal{O} \rightarrow \mathbb{R}^{s \times (d+1)}$  is a smooth function such that

<span id="page-16-1"></span>
$$
D_UQ_j(U)=\mathfrak{B}(U)D_UG_j(U), \text{ for all } U\in\mathcal{O}, j\in 0,\cdots,k, \quad (5)
$$

for some smooth function  $\mathfrak{B}:\mathcal{O}\to \mathbb{M}^{s\times n}.$  The function  $Q$  is called a *companion* of  $G$  and equation  $(4)$  is called a *companion* law of the conservation law  $(3)$ .

# How much regularity of a weak solution is required so that it also satisfies the companion law?

#### Theorem

Let  $U \in B_3^{\alpha,\infty}$  $\frac{\partial \alpha}{\partial 3}(\mathcal{X};\mathcal{O})$  be a weak solution of [\(3\)](#page-15-0) with  $\alpha>\frac{1}{3}$ . Assume that  $G \in \mathrm{C}^2(\mathcal{O};\mathbb{M}^{n\times (d+1)})$  is endowed with a companion law with flux  $Q \in C(\mathcal{O}; \mathbb{M}^{s \times (d+1)})$  for which there exists  $\mathfrak{B}\in \mathrm{C}^1(\mathcal{O};\mathbb{M}^{s\times n})$  related through identity  $(5)$  and the essential image of U is compact in  $\mathcal{O}$ . Then U is a weak solution of the companion law [\(4\)](#page-16-0) with the flux Q.

- 
- P. Gwiazda, M. Michálek and A. Świerczewska-Gwiazda,

A note on weak solutions of conservation laws and energy/entropy conservation. arxiv.1706.10154, 2017.

- **•** the generality of the above theorem is achieved at the expense of optimality of the assumptions.
- However given additional information on the structure of the problem at hand one might be able to relax some of these assumptions.
- the theorem provides for instance a conservation of energy result for the system of polyconvex elastodynamics, compressible hydrodynamics et al.
- **O T**. Debiec, P. Gwiazda, and A. Świerczewska-Gwiazda, A tribute to conservation of energy for weak solutions arXiv:1707.09794, 2017.

# <span id="page-19-0"></span>Thank you for your attention